Exploring Nature's Accelerators with the Fermi Gamma-ray Space Telescope



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The Sky at Night



The Sky at Night



Milky Way – Gamma rayshe from powerful cosmic ray particles smashing into the tenuous gas between the stars.

> Particle-Particle Collisions

and a second second

Pulsars – rapidly spinning neutron stars with enormous magnetic and electric fields

MPIfR-Bonn Pulsar Group

3

Blazars – AGN with jets pointed right at Earth.



Blazars – AGN with jets pointed right at Earth.

Gamma-ray bursts – extreme exploding stars or merging black holes or neutron stars.

The Unknown – many sources remain mysterious

Gamma-ray Astrophysics



- Gamma-rays cover a huge swath of the electromagnetic spectrum
- High Energy gamma-rays produced in non-thermal processes -> i.e. by energetic charged particles
 - Explore extreme environments hosting powerful particle accelerators
 - Natural connections to neutrino and cosmic-ray astrophysics

Fermi Observatory



Large Area Telescope (LAT):

- 20 MeV >300 GeV
- 2.4 sr FoV (scans entire sky every ~3hrs)

Gamma-ray Burst Monitor (GBM)

- 8 keV 40 MeV
- views entire unocculted sky

Launched on June 11, 2008

Detecting Gamma-rays

6

E=mc² matters!

- E Energy
- m mass
- c speed of light in a vacuum

Detecting Gamma-rays

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Large Area Telescope







Si Tracker pitch = 228 μm 8.8 10⁵ channels 18 planes

Y Large Area Telescope

Csl Calorimeter hodoscopic array (8 layers) 6.1 10³ channels

LAT: 4 x 4 modular array 3000 kg, 650 W 20 MeV – 300 GeV



The Sky Above 1 GeV

- >5000 sources of gamma-rays

The Sky above 50 GeV

- >250 sources
- With more observations we continue to push the upper energy range of Fermi-LAT higher



The Variable Gamma-ray Sky



Gamma-ray Burst Monitor



Detects flashes of gamma-rays at lower energies than the LAT

Bursts of Gamma-rays from Near and Far



rays from Near and Far





Sept. 1, 2014





Selected Results from the last 10 years

2008 – Solving a mystery and opening a new field

Abdo et al., 2008, Science



P ~ 316 ms Pdot ~ 3.6 x 10^{-13} Flux (>100MeV) = 3.8 \pm 0.2 x 10^{-7} ph cm⁻² s⁻¹

Pulse undetected in radio



35 Gamma-ray selected pulsars

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40 millisecond pulsars

42 Radio loud pulsars Now have over 200 Fermi-detected pulsars

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7 gamma-ray pulsars before Fermi (6 from EGRET and one from AGILE)

2009: Antimatter from Thunderstorms



18





2010: Fermi Bubbles

Bubble spectrum

Spectrum of diffuse gamma-ray emission

Brightness

Brightness

Bubble emission Diffuse emission

Gamma-ray emissions

2010: Fermi Bubbles

X-Ray emissions

Milky Way

50,000 light years

22

Sun

2011 – A not so standard candle





- Shockingly bright flares from the Crab Nebula in Sept 2010 and April 2011 observed by AGILE and Fermi-LAT
- Rapid (hourscale) variability of PeV electrons poses severe challenges for acceleration mechanisms

2012: A High Energy Solar Flare



- March 7, 2012 Fermi-LAT detected an intense solar flare lasting 20 hours with emission up to energies of 4 GeV
 - Highest energy photons ever seen from the Sun
 - Longest duration gamma-ray flare
 - First time localizing high energy gamma-ray emission on the Sun's disk

2013: Dark Matter At the Galactic Center?

Fermi detected an unexpected source of gamma-rays from the center of our Galaxy

Likely the densest dark matter source in the Galaxy

- Fermi observations show that spectrum and spatial distribution consistent with
 - signal of dark matter,
 - a yet-to-be-identified population of dim gamma-ray emitting objects.



Starting in Dec 2013, we spent a year with observations focused on the Galactic Center

2014: Stellar Novae Produce High Energy Gamma-rays



Fermi-LAT Collaboration, Science 2014

- Serendipitous discovery of gamma-ray transient at the time and location of V407 Cyg – found something that we were not looking for!
- Later found many more...

2015: Gamma-ray periodicity in an Active Galaxy

Binary black hole?



2016: A Sharper View

- A major revamp of the data processing pipeline produces sharper images (right) and more gamma-rays (lower right) dramatically improving Fermi Large Area Telescope performance.
- Expanded Fermi-LAT coverage to higher energies allows the first census of the sky between 50 GeV and 2 TeV (below).







2016: Resolving a Proton Accelerator

- Improved data resolve the shell of supernova remnant IC 443 at physical scales of ~5 pc
- Dense molecular and diffuse atomic (fast shock) regions differ in brightness by ~10x, but spectra are surprisingly consistent





2017: Gamma-ray Burst and Gravitational Waves





Time from merger (seconds)





The first 10 years of Fermi have been great – I'm looking forward to what is yet to come!

2018: Neutrinos and gamma-rays







2009 – Testing Einstein's Theory with a Gamma-ray Burst





- Highest energy gamma-ray arrives within 0.9s of the lower energy photons after traveling 7 billion years
- Eliminates theories of quantum gravity which predict that space-time is "foamy" enough to interfere strongly with light.

Gamma-ray Burst Monitor



Detects flashes of gamma-rays at lower energies than the LAT

2011: Cosmic Fireworks!



2015: Record Flare from the Active Galaxy 3C279



 Huge rapid flare from 3C279 challenges standard ideas of particle acceleration in relativistic jets



2018: Stay Tuned for Another Fermi Multimessenger Announcement on July 12!

