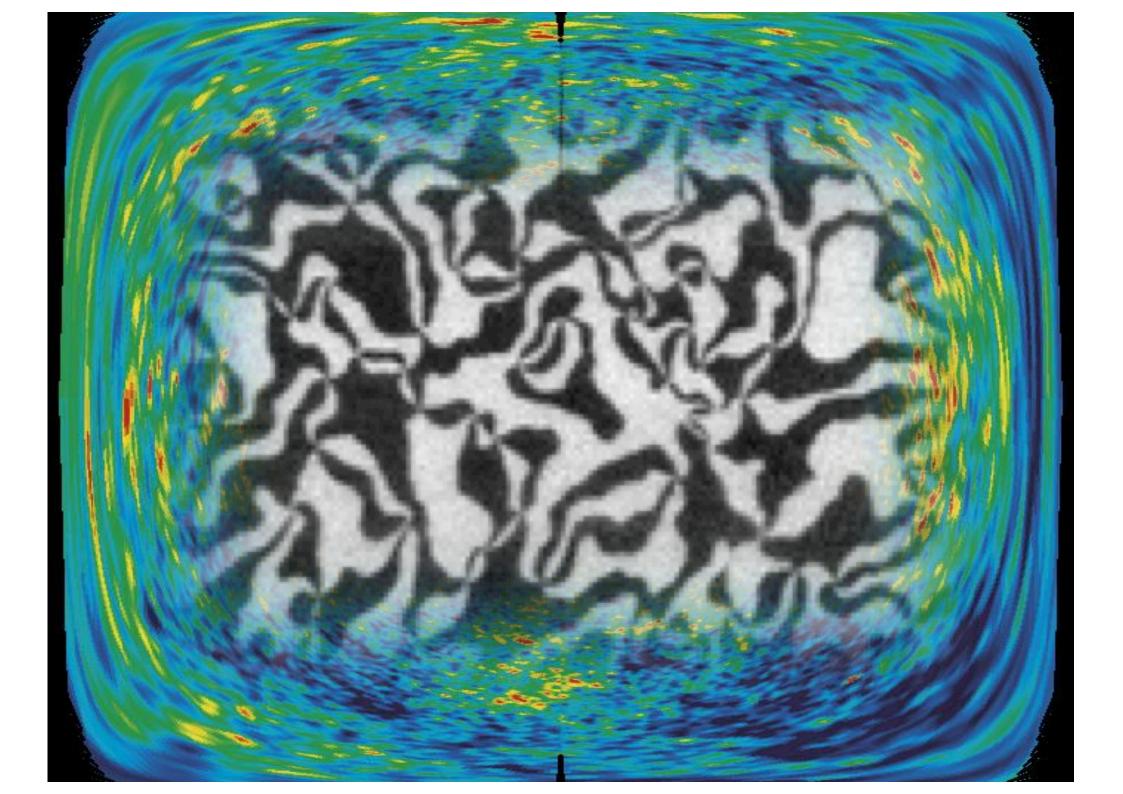
MATERIALS THEORY

From materials to cosmology Studying the early universe under the microscope





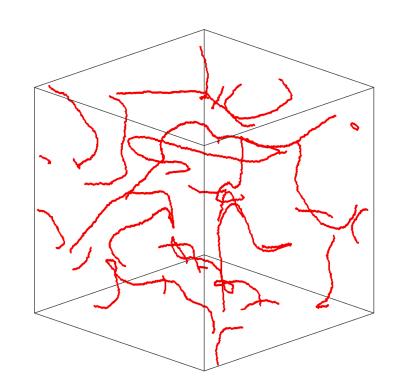
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MATERIALS THEORY







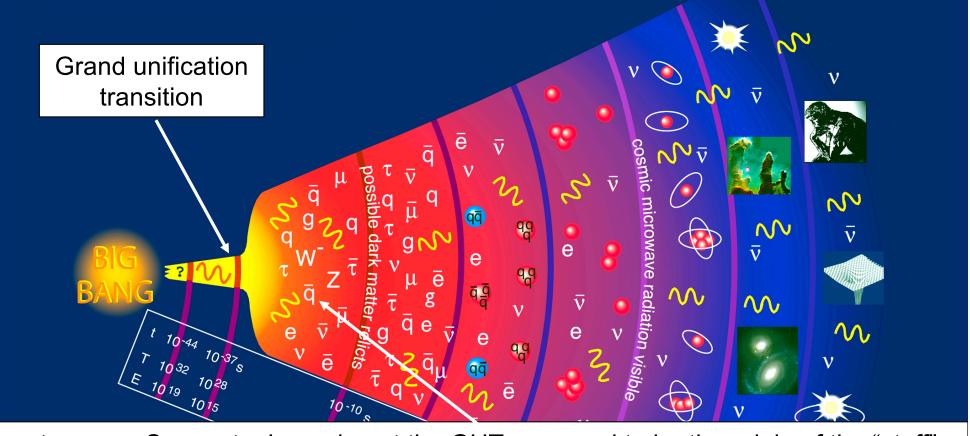
N.A. Spaldin, *Multiferroics, from the cosmically large to the subatomically small*, Nat. Rev. Mater. 2, 17017 (2017)

 $\hat{H}\Psi(\mathbf{r},t)$

"Standard Model" of Cosmology



"Standard Model" of Cosmology



Spontaneous Symmetry Lowering at the GUT proposed to be the origin of the "stuff"

10-0

10 -1

Particle Data Group, LBNL, © 2008. Supported by DOE and NSF (GeV)

 $\overline{\mathbf{v}}$

109

10-12

3x105

3000

3x10-10

n

n

Today

12x109y (sec, yrs)

ETH

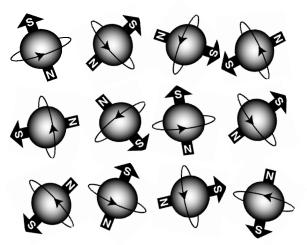
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MATERIALS THEORY

Spontaneous symmetry lowering phase transition in a ferromagnet

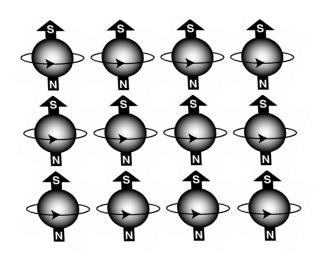
 $H\Psi(\mathbf{r},t)$

high temperature



high symmetry

Transition (Curie) temperature, Tc



low symmetry

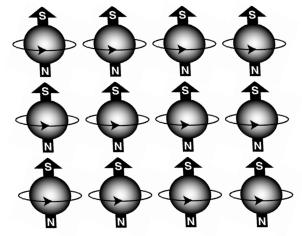
low temperature

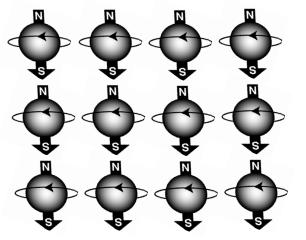
MATERIALS THEORY

1. Choice of multiple equivalent low-symmetry states

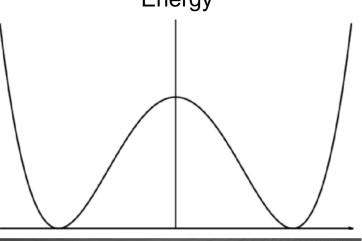
e.g. in a uniaxial ferromagnet there are two equivalent low symmetry states

 $\hat{H}\Psi(\mathbf{r}, \mathbf{r})$





and the symmetry-lowering phase transition is described by a double well potential: Energy

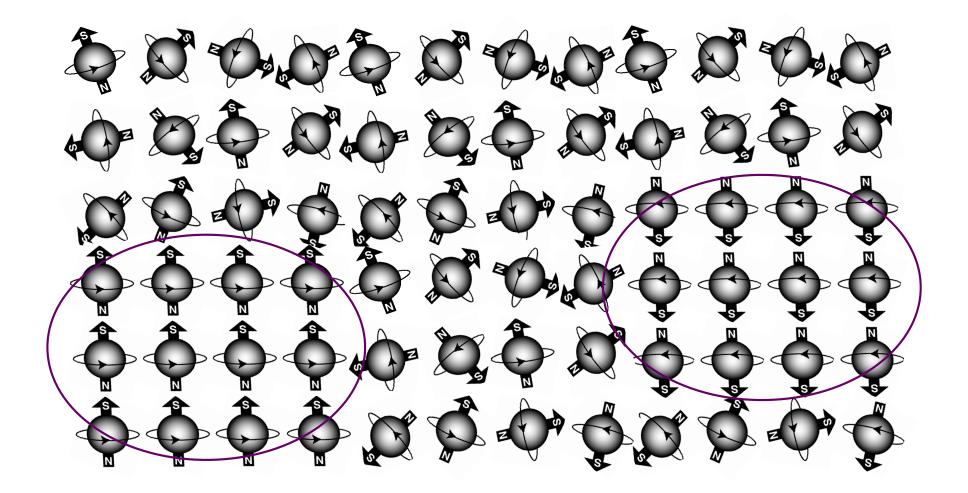


Order parameter / "line-up-ness" of spins



2. Defect formation at symmetry-lowering phase transitions

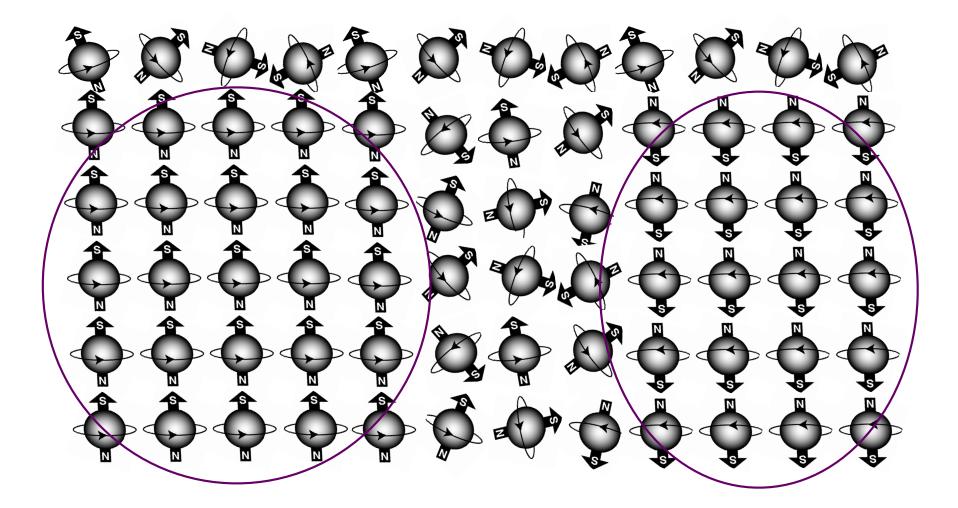
 $H\Psi(\mathbf{r},t)$





2. Defect formation at symmetry-lowering phase transitions

 $\hat{H}\Psi(\mathbf{r},t)$

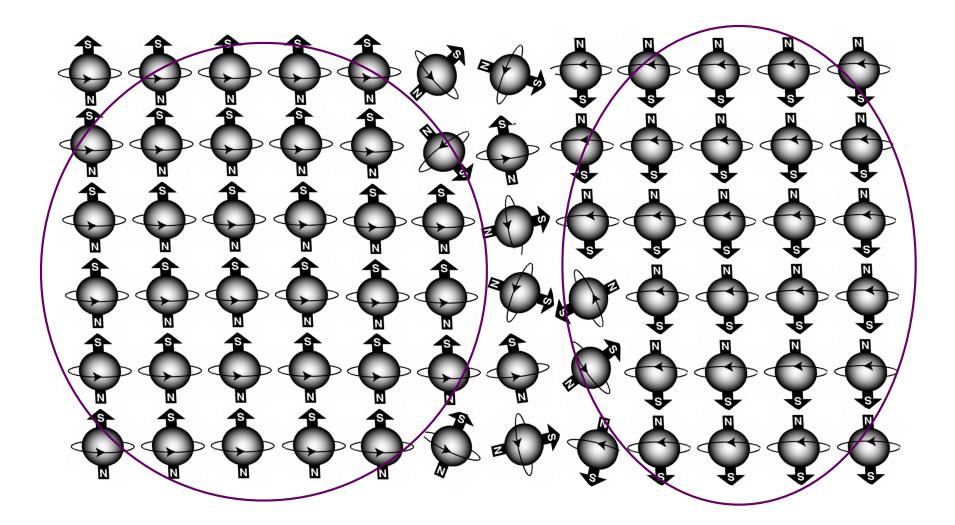




2. Defect formation at symmetry-lowering phase transitions

 $\hat{H}\Psi(\mathbf{r},t)$

 ∂t

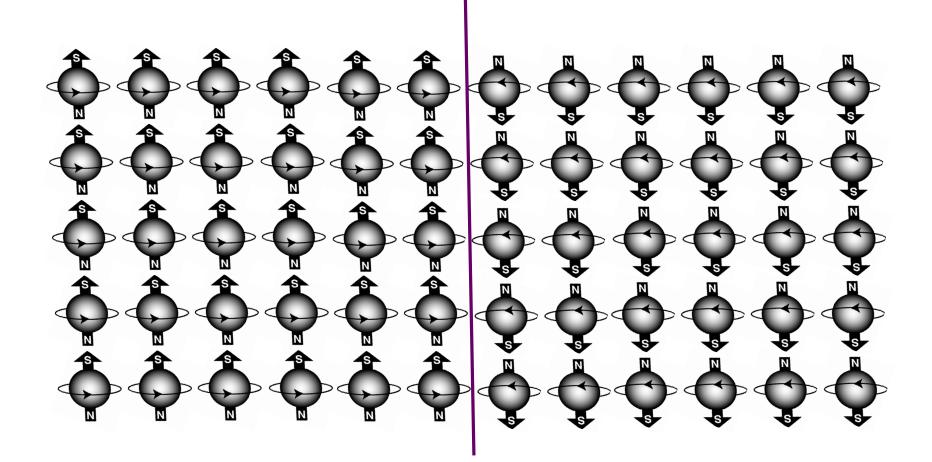




2. Defect formation at symmetry-lowering phase transitions

 $\hat{H}\Psi(\mathbf{r},t)$

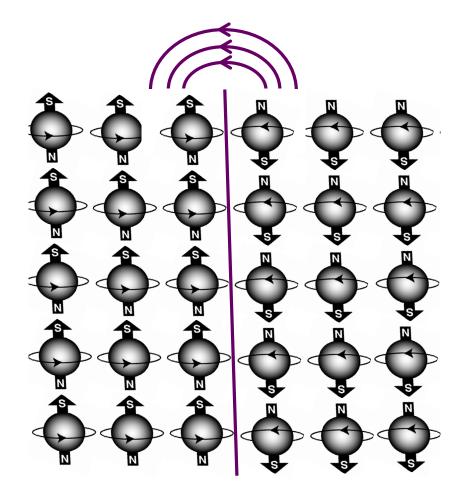
 ∂t

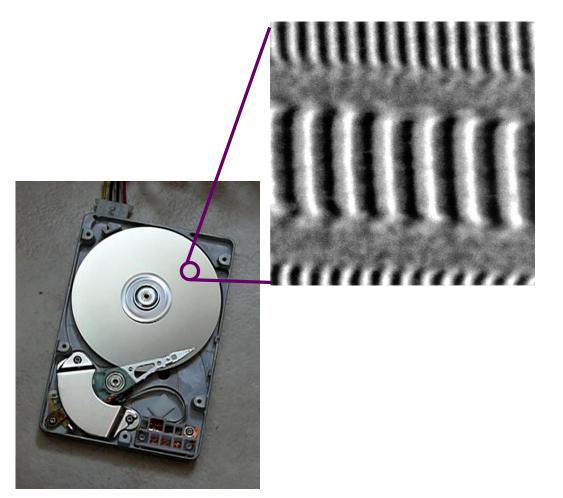




MATERIALS THEORY

Defects in ferromagnets: Domain walls





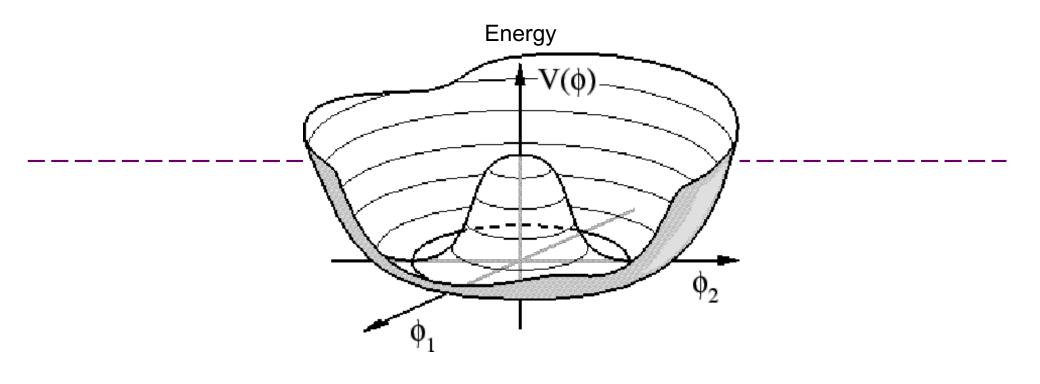
 $H\Psi(\mathbf{r},\mathbf{r})$

Materials Theory

Spontaneous symmetry lowering at the Grand Unification Transition

HOMOGENEOUS VACUUM

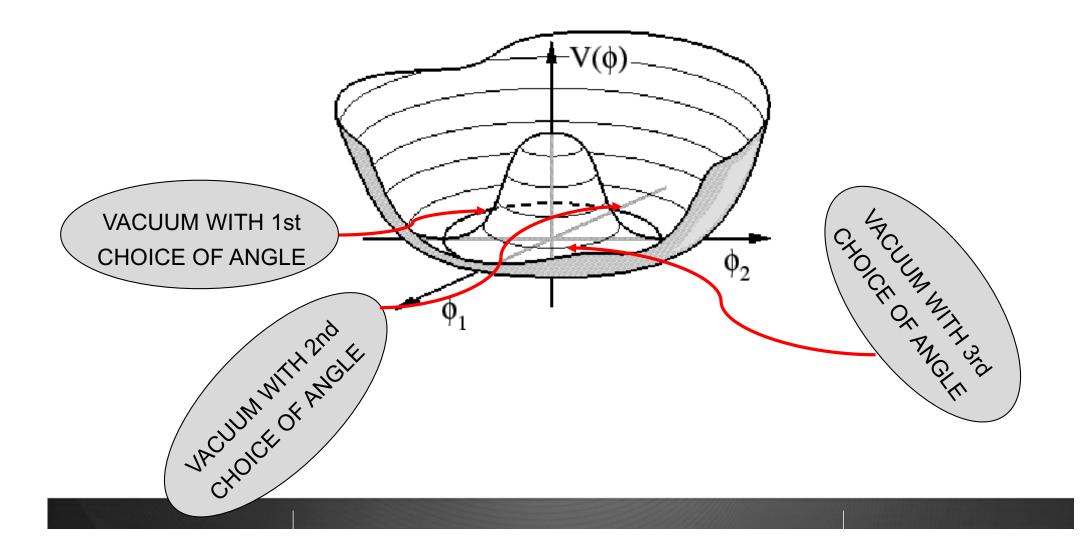
 $\hat{H}\Psi(\mathbf{r},t)$



LOWER SYMMETRY VACUUM

Materials Theory

Different regions of the early universe choose different equivalent ground states

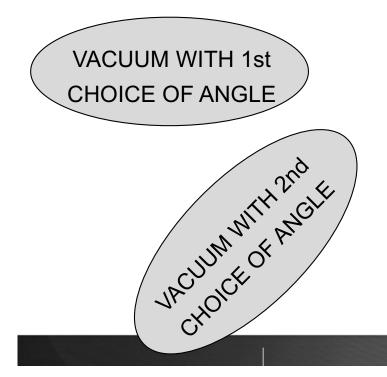


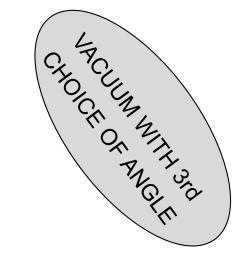


 $H\Psi(\mathbf{r}, \mathbf{r})$

 $\hat{H}\Psi(\mathbf{r},t)$

Materials Theory

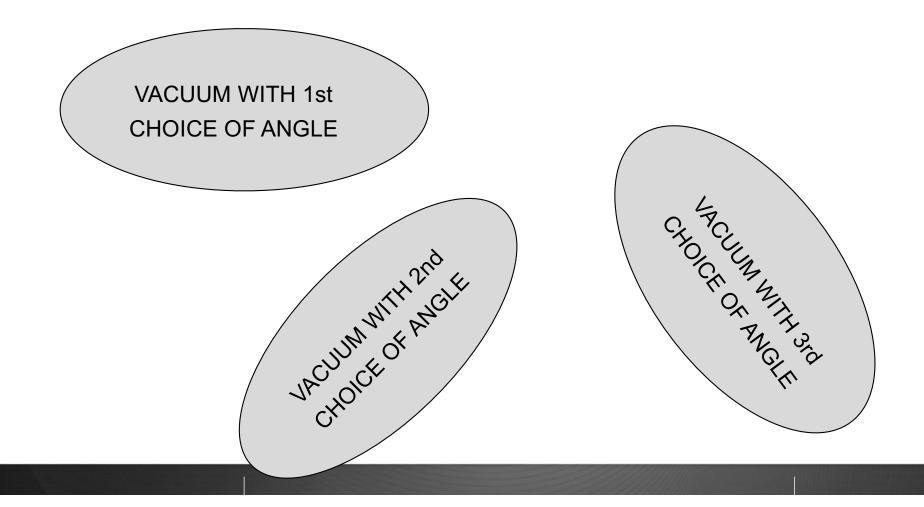




MATERIALS THEORY

As the universe expands through the transition, the low symmetry regions grow...

 $\hat{H}\Psi(\mathbf{r},\mathbf{r})$

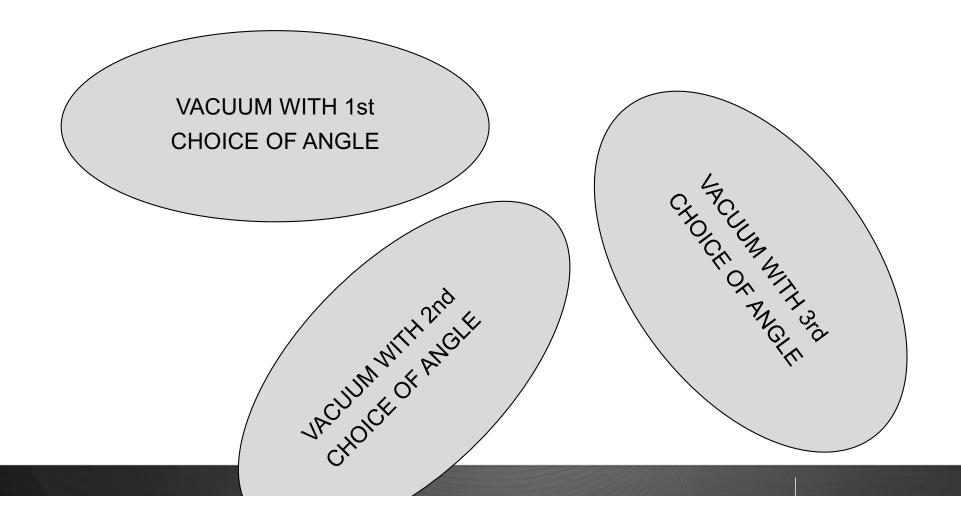


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Eidgenössische Technische Hochschule Zürich Swiss Federal Institute of Technology Zurich $H\Psi(\mathbf{r})$

MATERIALS THEORY

and grow...



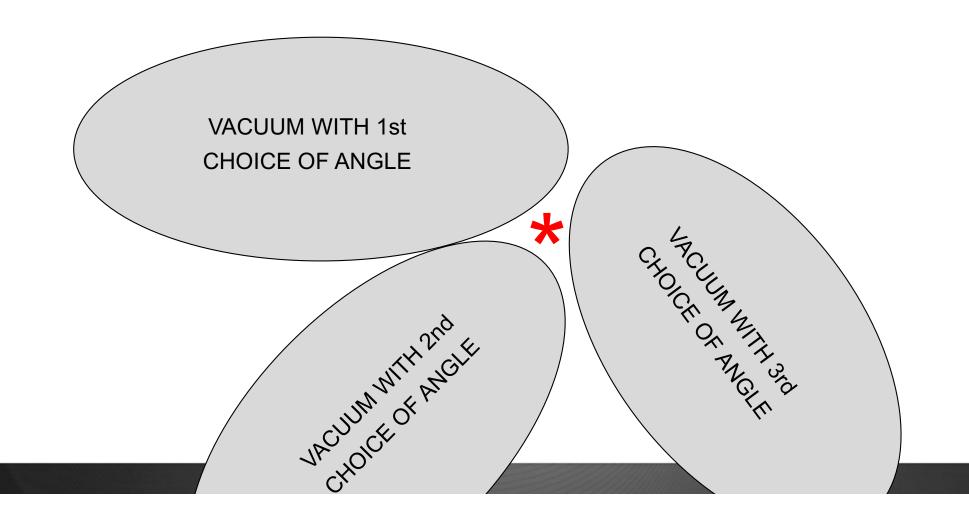
 $\hat{H}\Psi(\mathbf{r},t)$

ETH

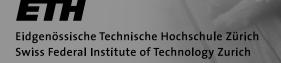
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MATERIALS THEORY

and eventually meet!

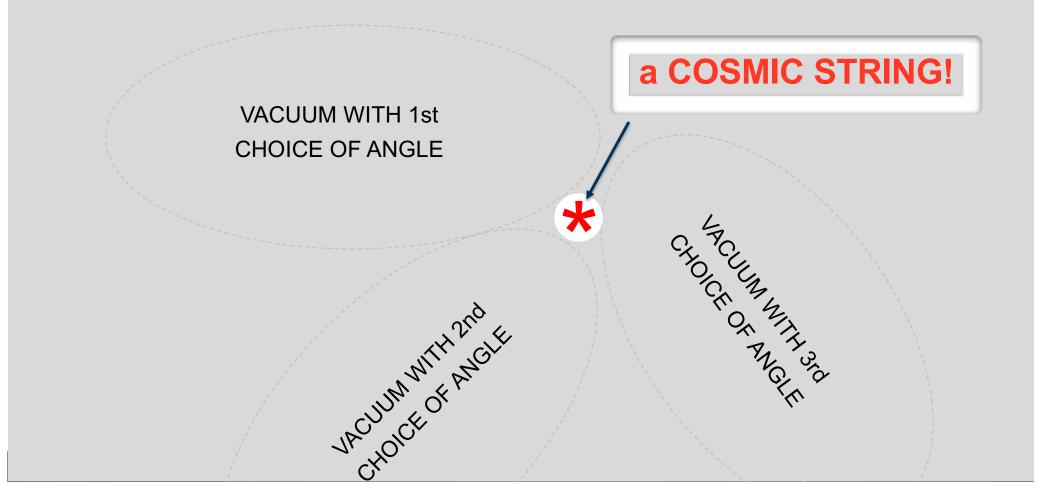


 $\hat{H}\Psi(\mathbf{r},t)$



The angle mismatch in the low-symmetry vacuum is a topologically protected one-dimensional defect (Kibble)

 $\hat{H}\Psi(\mathbf{r})$



A detail: How many cosmic strings should we have?

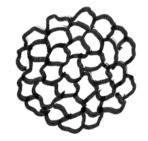
It depends on the rate of expansion of the early universe (Zurek)

Expand slowly: Different regions can communicate their choice of angle

- \rightarrow Large regions of the same choice
- \rightarrow Low density of cosmic strings

Expand quickly: Not much time to communicate choice of angle

- \rightarrow Many smaller regions with different choices of angle
- \rightarrow High density of cosmic strings





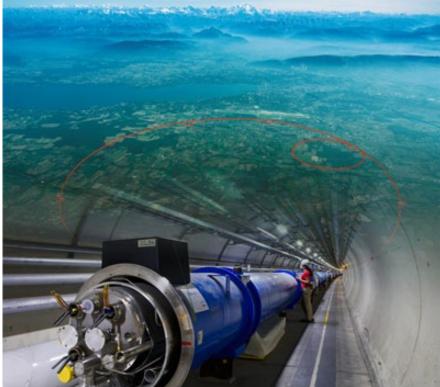
Materials Theory

Do cosmic strings exist? How can we study them?

 $H \mathbb{U}(\mathbf{r})$

For direct study we need a probe with a similar energy, $\sim 10^{15}$ GeV

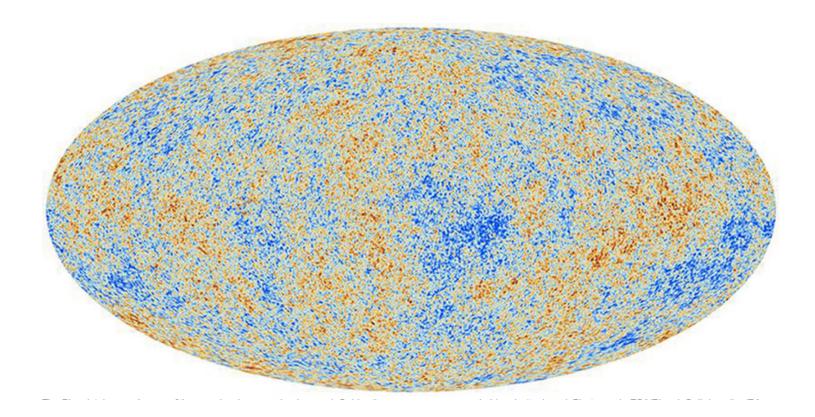
Our highest energy probes, the largest hadronic colliders reach ~10,000 GeV



Materials Theory

How is Cosmic String Formation at the Grand Unification Transition studied?

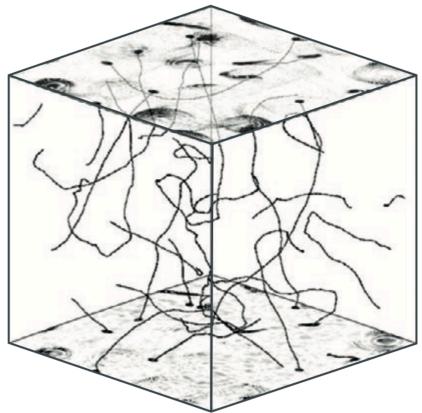
Analyzing the Cosmic Microwave Background



Materials Theory

How is Cosmic String Formation at the Grand Unification Transition studied?

Computer Simulation



PHYSICAL REVIEW D 76, 043005 (2007)

CMB polarization power spectra contributions from a network of cosmic strings

Neil Bevis,^{1,*} Mark Hindmarsh,^{1,†} Martin Kunz,^{2,‡} and Jon Urrestilla^{1,§}

Instead we will study the GUT in our laboratory!

Outline

First we will identify a material with a symmetry-lowering phase transition described by the same mathematics as that proposed for the GUT

spontaneous symmetry breaking described by a Mexican hat potential

 $\hat{H}\Psi(\mathbf{r},t)$

Then we will do experiments on the material to answer questions about the GUT:

> Do cosmic strings exist? Did they form as we think? How did they evolve? What are their properties?

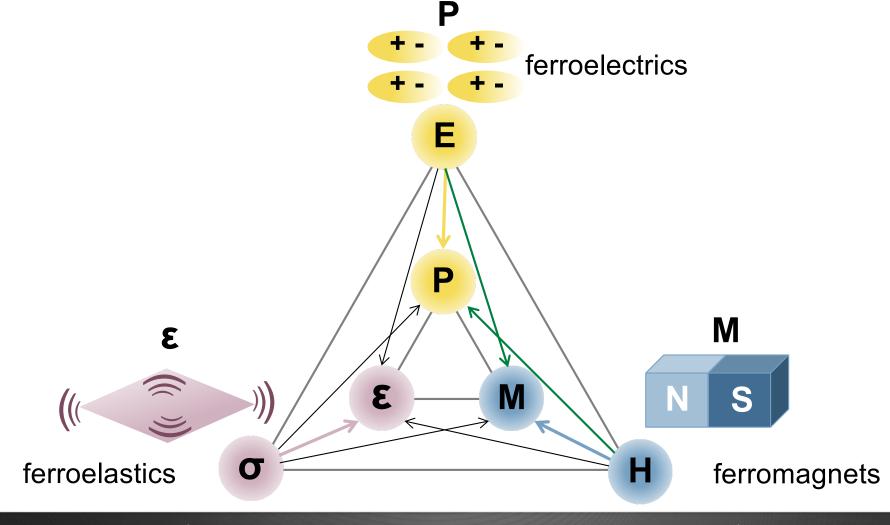


Sinead Griffin



MATERIALS THEORY

Where to find a suitable material? Multiferroics: Multiple ferroic orders...



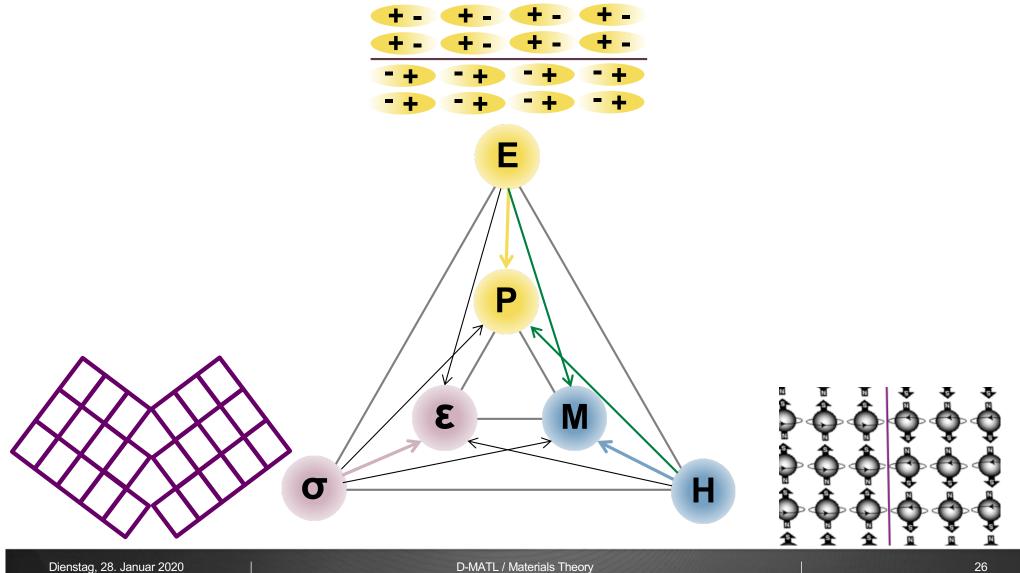
 $\hat{H}\Psi(\mathbf{r},t)$



MATERIALS THEORY

...and multiple defects from spontaneous symmetry-lowering transitions

 $\hat{H}\Psi(\mathbf{r},t)$

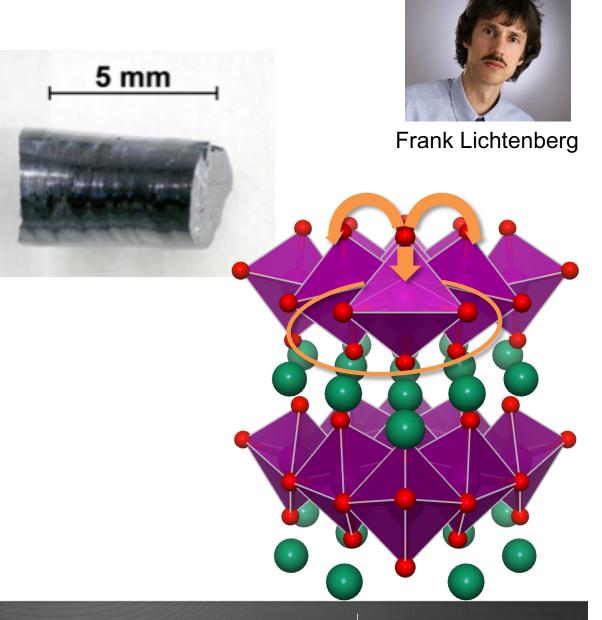




MATERIALS THEORY

Our material: Multiferroic YMnO₃

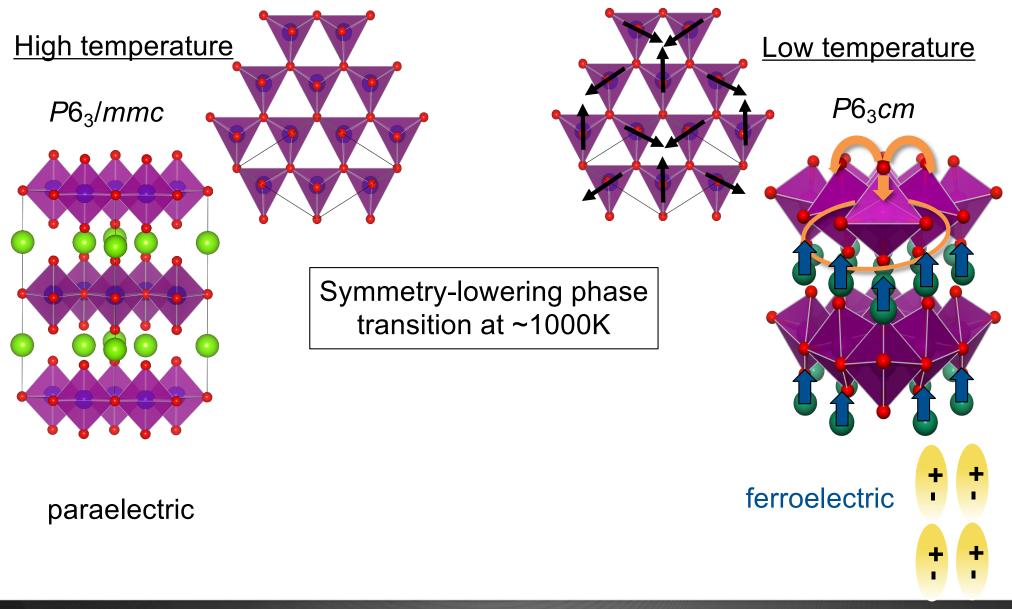






MATERIALS THEORY

Our material: Multiferroic YMnO₃

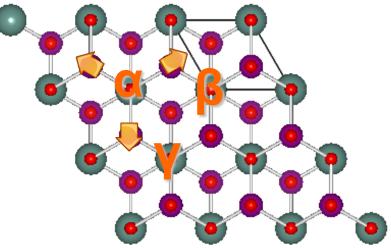




MATERIALS THEORY

Look at these distortions in more detail:

Trimerization / Tilting



Three possible origins





MATERIALS THEORY

Look at these distortions in more detail:

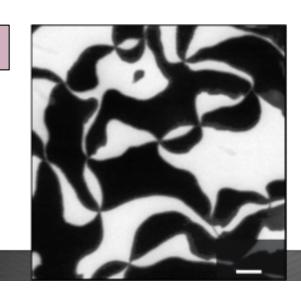
Trimerization / Tilting

Three possible origins

Polarization +P

-P

Two possible orientations



 $H\Psi(\mathbf{r},$

Results in six domains

Materials Theory

Calculate the form of the potential using symmetry analysis and density functional theory

 $H\Psi(\mathbf{r},t)$

Landau free energy

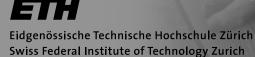
$$f_{\rm u} = \frac{a}{2}Q^2 + \frac{b}{4}Q^4 + \frac{Q^6}{6}\left(c + c'\cos 6\Phi\right) - gQ^3P_z\cos 3\Phi$$

Q is amplitude of tilting

- Φ is angle of tilting
- P_z is polarization

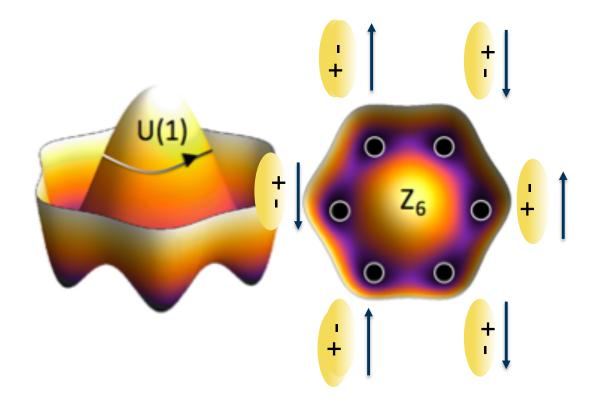
The phase transition is described by a Mexican-hat-like potential!

S. Artyukhin, K.T. Delaney, NAS and M. Mostovoy, *Landau theory of topological defects in multiferroic hexagonal manganites*, Nature Materials, 13, 42 (2014)



And the details of our "Mexican Hat-like" potential make it easy for us to measure!

 $\hat{H}\Psi(\mathbf{r},t)$



S. Artyukhin, K.T. Delaney, NAS and M. Mostovoy, *Landau theory of topological defects in multiferroic hexagonal manganites*, Nature Materials, 13, 42 (2014)



MATERIALS THEORY

And the details of our "Mexican Hat-like" potential make it easy for us to measure!

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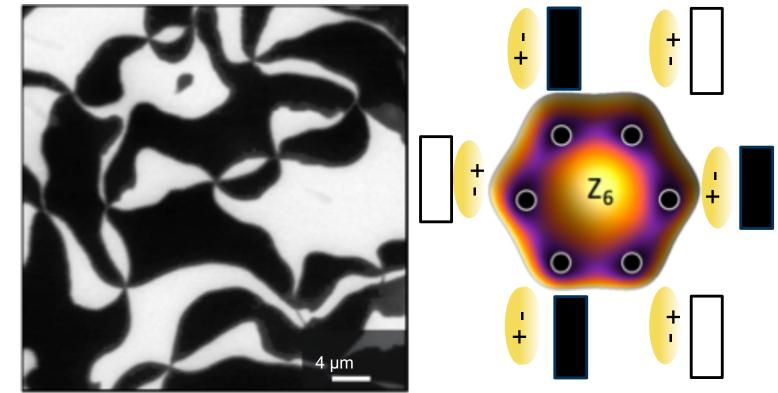
Martin Lilienblum





Manfred Fiebig

<u>Piezoforce Microscopy Image of</u> <u>ferroelectricity in YMnO₃</u>



S. Artyukhin, K.T. Delaney, NAS and M. Mostovoy, *Landau theory of topological defects in multiferroic hexagonal manganites*, Nature Materials, 13, 42 (2014)

D-MATL / Materials Theory

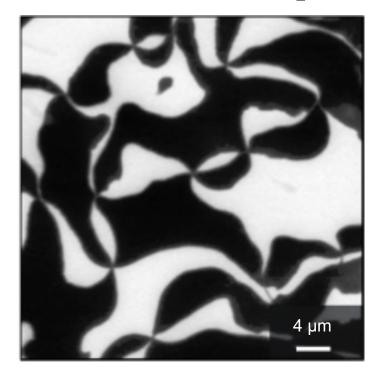
ETH

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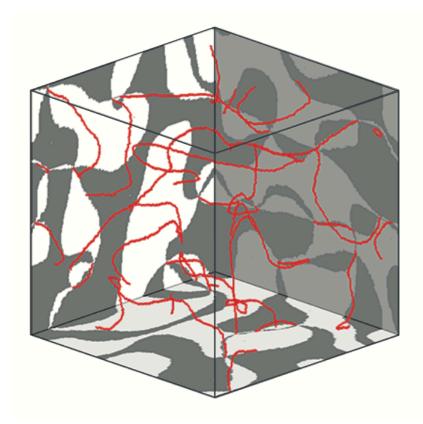
MATERIALS THEORY

The meeting points of the ferroelectric domains are in fact one-dimensional "strings"

Piezoforce Microscopy Image of the Defects in YMnO₃



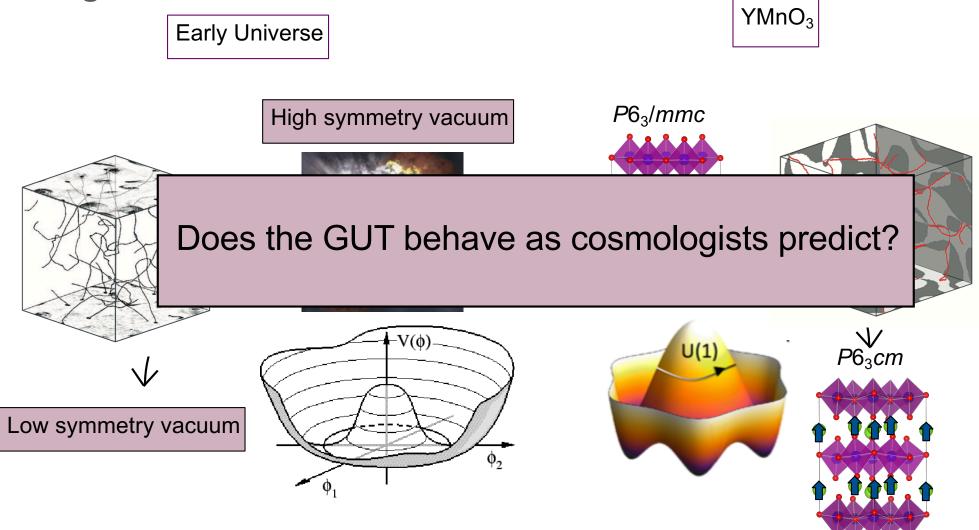
3-D Simulation



Thomas Lottermoser, Fiebig group

The structural phase transition in multiferroic YMnO₃ provides an analogue to the Grand Unification Transition

 $\hat{H}\Psi(\mathbf{r},t)$



N.A. Spaldin, Multiferroics, from the cosmically large to the subatomically small, Nat. Rev. Mater. 2, 17017 (2017)

What experiment would we like to do on the early universe?

We'd like to expand it at different rates, crossing the GUT, and see how many cosmic strings form in each case ("Kibble-Zurek scaling")

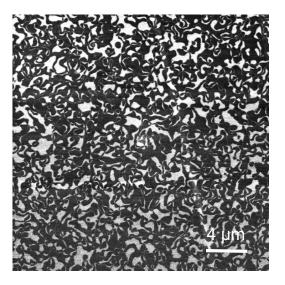
Expand slowly: Different regions can communicate their choice of angle

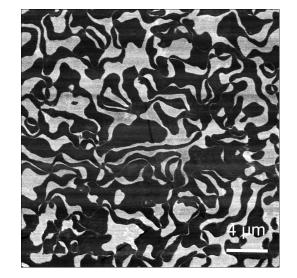
- \rightarrow Large regions of the same choice
- \rightarrow Low density of cosmic strings

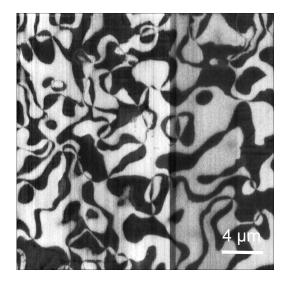
Expand quickly: Not much time to communicate choice of angle

- \rightarrow Many smaller regions with different choices of angle
- \rightarrow High density of cosmic strings

Instead we will cool YMnO₃ at different rates through the structural phase transition and count how many "strings" form







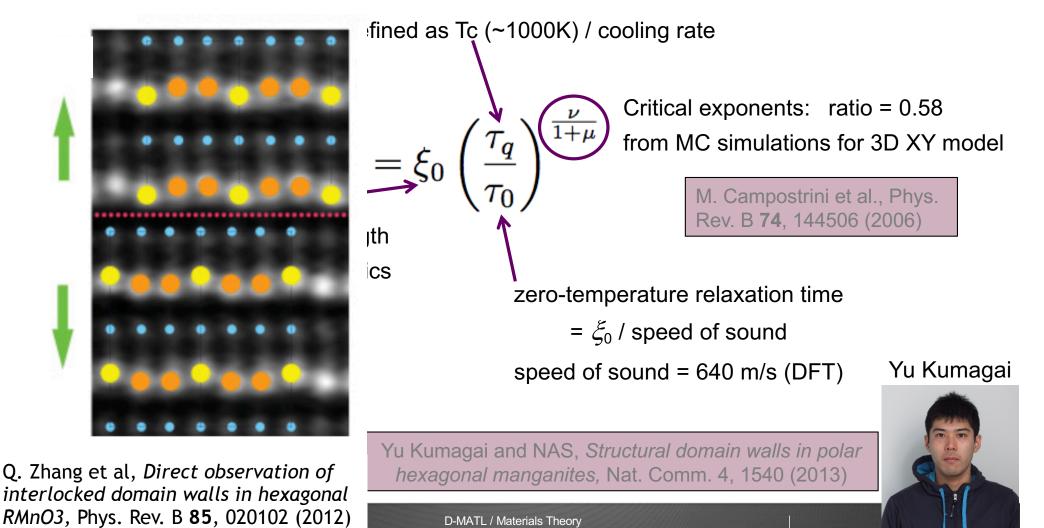




Testing the predicted "Kibble-Zurek scaling" for defect formation

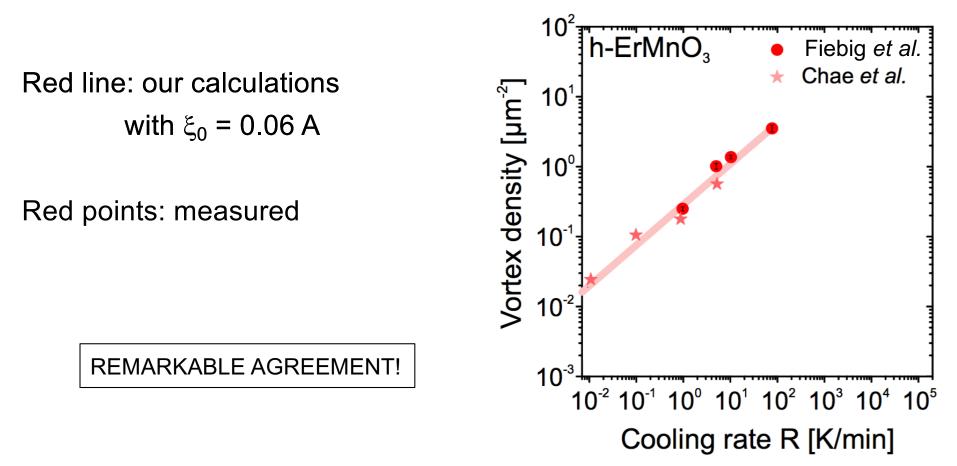
 $H\Psi(\mathbf{r},t)$

domain size (for linear quench),



Comparison of predicted Kibble-Zurek scaling with experiment

 $\hat{H}\Psi(\mathbf{r},t)$

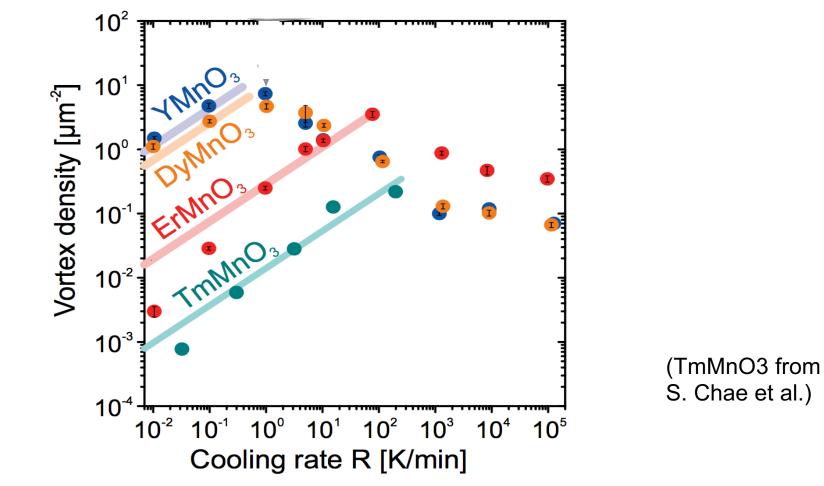


S. C., Chae et al., Direct observation of the proliferation of ferroelectric loop domains and vortex-antivortex pairs, PRL **108**, 167603 (2012) S. Griffin, M. Lilienblum, K. Delaney, Y. Kumagai, M. Fiebig and N. A. Spaldin, Scaling behaviour and beyond

equilibrium in the hexagonal manganites, PRX 2, 041022 (2012)



"Beyond-KZ regime" and unexpected dependence on chemistry

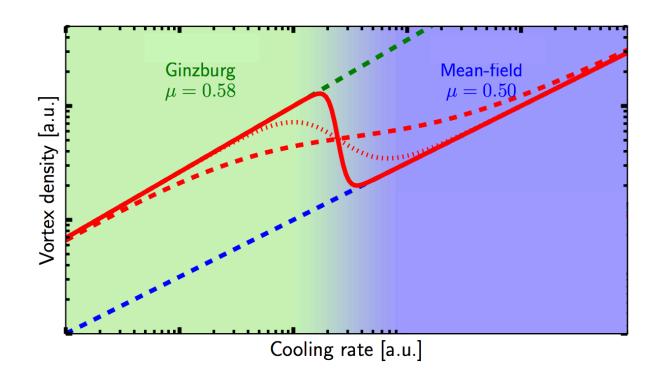


PHYSICAL REVIEW X 7, 041014 (2017) Global Formation of Topological Defects in the Multiferroic Hexagonal Manganites

Q. N. Meier,¹ M. Lilienblum,¹ S. M. Griffin,² K. Conder,³ E. Pomjakushina,³ Z. Yan,^{4,5} E. Bourret,⁴ D. Meier,⁶ F. Lichtenberg,¹ E. K. H. Salje,⁷ N. A. Spaldin,¹ M. Fiebig,¹ and A. Cano^{1,8}

Open questions:

What causes the chemistry dependence? What is the origin of the turnaround? What is the physics of the beyond-KZ regime?





Quintin Meier

Is it relevant for early-universe behavior?

HU(r

The

Summary

Multiferroics provide the first example of Kibble-Zurek scaling in a solidstate system

Cosmic strings formed the way cosmologists think ;)

Table-top astrophysics How to build a multiverse Economist

 $\hat{H}\Psi(\mathbf{r},t)$

Small models of cosmic phenomena are shedding light on the real thing

Mar 16th 2013 From the print edition

f Like

Whether all this ingenuity unravels any cosmic truth is uncertain. Cliff Burgess, a theorist at Perimeter Institute for Theoretical Physics in Ontario, has his doubts. But he thinks that such experiments are nevertheless worth pursuing. "Like tap-dancing snakes," he says, "the point is not that they do it well, it is that they do it at all."

